

The Hawking–Penrose Singularity Theorem for C^1 -Lorentzian Metrics

Michael Kunzinger, Argam Ohanyan*, Benedict Schinnerl, Roland Steinbauer

Department of Mathematics, University of Vienna

12th Central European Relativity Seminar
Hungarian Academy of Sciences, Budapest, Hungary
February 21st, 2022

Singularity Theorems: General Remarks

Yield existence of incomplete causal geodesics under reasonable assumpt.

Theorem (Pattern singularity theorem, [Senovilla, 98])

A spacetime has incomplete causal geodesics if it satisfies

(E) an energy (i.e. curvature) condition,

(C) a causality condition,

(I) an initial/boundary condition.

"Proof":

(I) → initially, geodesics start focusing.

(E) → they focus even more → focal points.

(C) → no focal points.

Conclusion: Not all causal geodesics can exist for all times.

Singularity Theorems: Matters of Regularity

- Classical singularity theorems valid for C^2 -spacetimes.
- Weak point: Extensions \leftarrow regularity?
- Goal: Obstruct complete low regularity extensions
 \rightarrow study singularity theorems for low reg. metrics.
- Ex.: Schwarzschild is C^0 -inextendible to $\{r = 0\}$ [Sbierski, 18].
- In many examples (e.g. matched spacetimes): metric is below C^2 .
- $g \in C^{1,1}$: Finite jumps in matter variables, bounded curvature, unique geodesics \rightarrow exp-map, convex/normal neighborhoods;
 \nexists notion of conjugate/focal points.
(Ref.: Hawking theorem [KSSV15], Penrose theorem [KSV15],
Hawking–Penrose theorem [GGKS18]).
- Recent new step: $g \in C^1$:
 - Hawking and Penrose singularity theorems [Graf, 20],
 - Gannon-Lee theorems [Schinnerl, Steinbauer, 21].

The Hawking–Penrose theorem for smooth spacetimes

Theorem ([Hawking, Penrose, 70])

Let (M, g) be a smooth spacetime satisfying the following:

- (i) (M, g) is chronological.
- (ii) $\text{Ric}(X, X) \geq 0$ for causal $X \in TM$ (SEC).
- (iii) Genericity holds along each inextendible causal geodesic, i.e. $\forall \gamma \exists t: R : [\gamma'(t)] \rightarrow [\gamma'(t)], v \mapsto [R(v, \gamma'(t))\gamma'(t)]$ is not identically zero.
- (iv) There exist one of the following:
 - (a) a compact, achronal, edgeless set;
 - (b) a trapped point;
 - (c) a trapped surface;

Then (M, g) contains incomplete causal geodesics.

(iv) with trapped submfds of arb. codimension [Galloway, Senovilla, 10].

From $C^{1,1}$ to C^1

- Geodesic eq. solvable, but not uniquely \rightarrow geodesic branching.
- \nexists exponential map, convex/normal neighborhoods.
- Geodesics are not locally maximizing.
- Levi-Civita connection $\Gamma \sim \partial g \in C^0$
 - \rightarrow curvature $\sim \partial \Gamma \in \mathcal{D}'^{(1)}$ (first order tensor distribution)
 - \rightarrow can insert C^1 -vector fields into curvature tensors
 - \rightarrow needed for distributional genericity condition
(loc. extending vector fields via parallel transport gives C^1 -fields):
Classically: $R(\cdot, \gamma'(t))\gamma'(t)$ not identically zero.
Distributionally: $g(R(V, X)X, V) > c \forall$ local C^1 -fields X, V close to $\gamma', \perp \gamma'$.

Hawking–Penrose in C^1 : methods of proof

- "Analytical regularization": $g \star_M \rho_\varepsilon \xrightarrow{C^1} g$ (\sim convolution).
- "Causal regularization": $\hat{g}_\varepsilon, \check{g}_\varepsilon$ with

$$\check{g}_\varepsilon \prec g \prec \hat{g}_\varepsilon \quad (\prec \dots \text{narrower lightcones})$$

and $\check{g}_\varepsilon, \hat{g}_\varepsilon \xrightarrow{C^1} g$ [Chrusciel, Grant, 12].

- $\hat{g}_\varepsilon, \check{g}_\varepsilon$ and $g \star_M \rho_\varepsilon$ are compatible:
Difference $\rightarrow 0$ in C_{loc}^∞ , difference of curvatures $\rightarrow 0$ in C_{loc}^0 , not just in $\mathcal{D}'^{(1)}$ [Graf, 20].
- Why both? E.g. energy conditions:
 - $\text{Ric}[g] \geq 0$ in $\mathcal{D}'^{(1)}$ $\rightarrow \text{Ric}[g] \star_M \rho_\varepsilon \geq 0$ in C^∞
 - $\text{Ric}[g] \star_M \rho_\varepsilon - \text{Ric}[g \star_M \rho_\varepsilon] \rightarrow 0$ in C^0
 - $\text{Ric}[g \star_M \rho_\varepsilon] - \text{Ric}[\check{g}_\varepsilon] \rightarrow 0$ in C_{loc}^0 \Rightarrow almost energy condition for \check{g}_ε .

The C^1 Hawking–Penrose singularity theorem

Theorem ([Kunzinger, O., Schinnerl, Steinbauer, 21])

Let (M, g) be a C^1 -spacetime satisfying the following:

- (i) (M, g) is causal.
- (ii) The **distributional** timelike and null energy conditions hold.
- (iii) **Distributional** genericity holds along each inextendible causal geodesic.
- (iv) (M, g) is **maximally causally nonbranching**.
- (v) There exist one of the following:
 - (a) a compact, achronal, edgeless set;
 - (b) a trapped point in the support sense;
 - (c) a trapped C^0 -surface in the support sense;
 - (d) a trapped C^0 -submanifold of codimension $2 < m < \dim M$ whose support submanifolds satisfy the **distributional** Galloway-Senovilla curvature condition.

Then (M, g) contains incomplete causal geodesics.



Outlook: How low can we go?

- *Causally*: Very low (cone structures, [Minguzzi, 19]).
- *Analytically*: Ultimate goal = Geroch-Traschen metrics.
Probably worth it: H^s -regularity (IVP).
- *Synthetically*: Lorentzian length spaces [KS, 17] (singularity theorems for LLS in [GKS18], [CM20]).
- *Compatibility*:
Causal \leftrightarrow analytic: OK.
Causal \leftrightarrow synthetic: OK.
Analytic \leftrightarrow synthetic: Unknown!

Selected references



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