

Interferometry of Entangled Quantum States of Light in Curved Space-Time

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Work done in collaboration with Christopher Hilweg.

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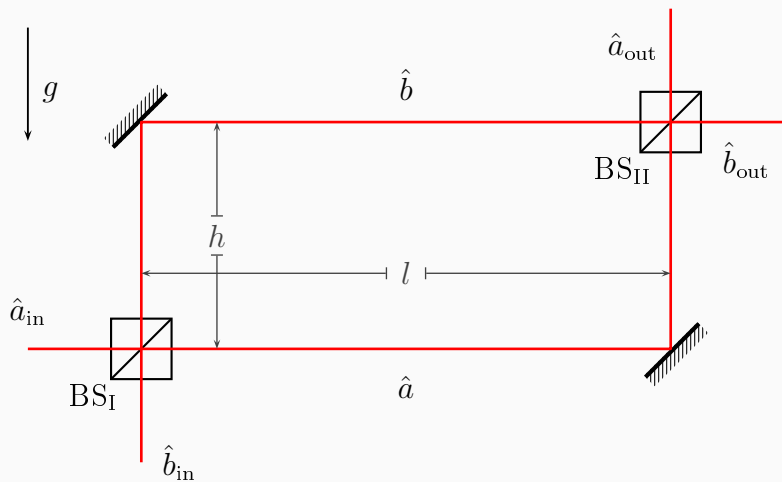
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Context: quantum physics in the presence of external gravitational fields

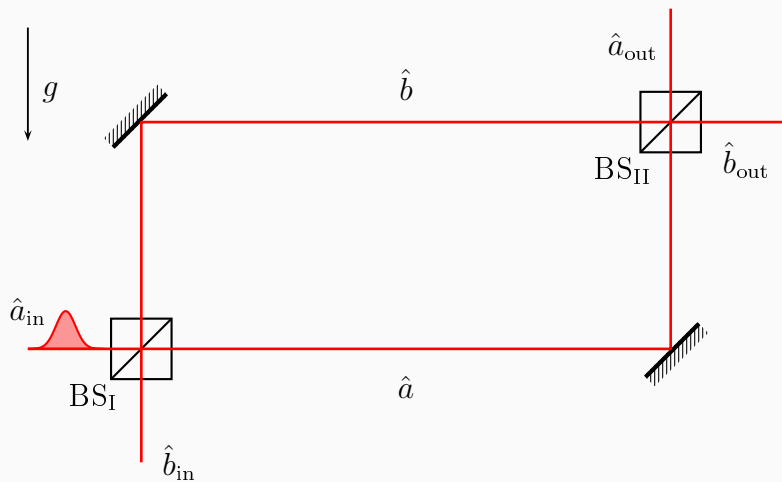
Ongoing efforts to demonstrate gravitationally-induced quantum interference

- Matter-wave interferometry (Colella *et al.* 1975; Kasevich & Chu 1992; Sorrentino *et al.* 2012) – **Newtonian gravity** suffices
- Optical interferometry (Zych *et al.* 2012; Hilweg *et al.* 2017) – **beyond Newtonian description**, but insensitive to **gravity gradients**

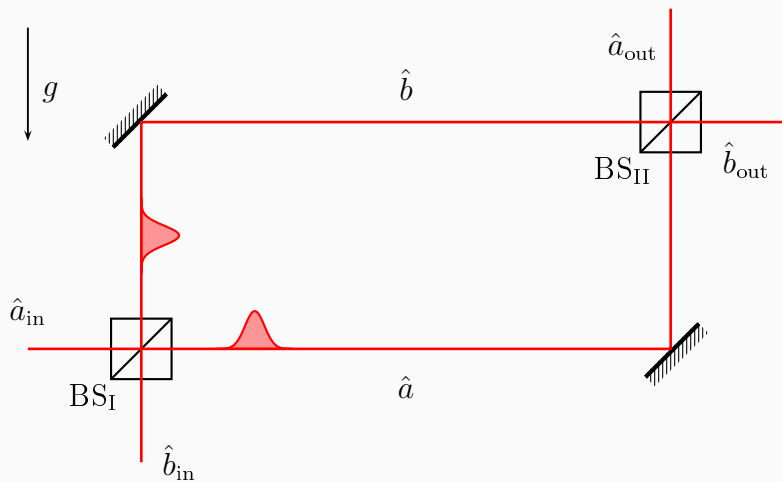
So far: no experimental realisation of quantum interference, where the phase shift is explicable only using general relativity



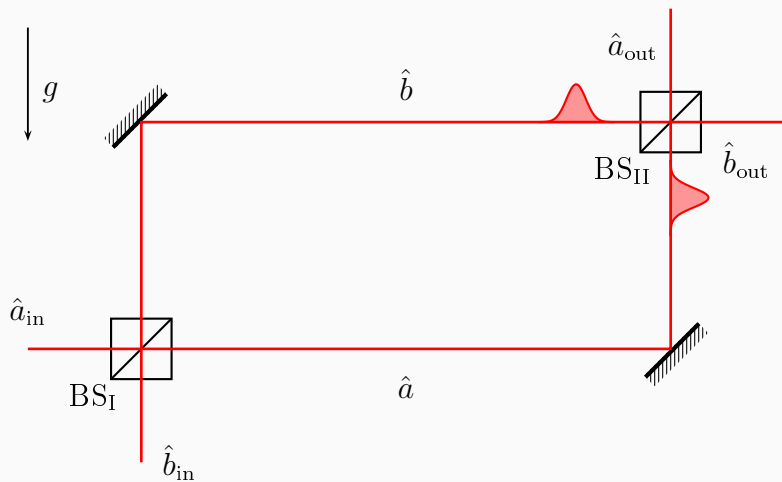
Single Photons



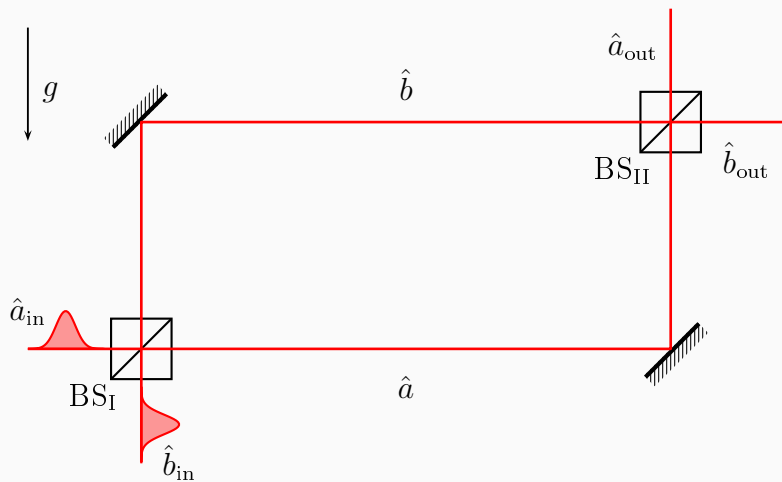
Single Photons



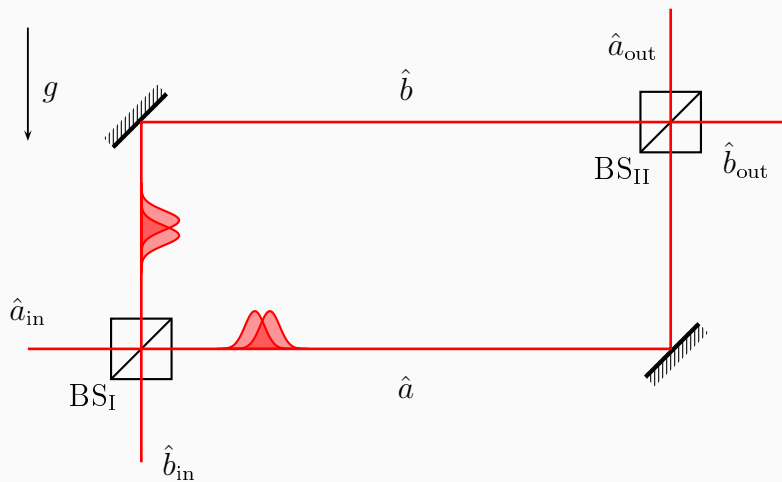
Single Photons



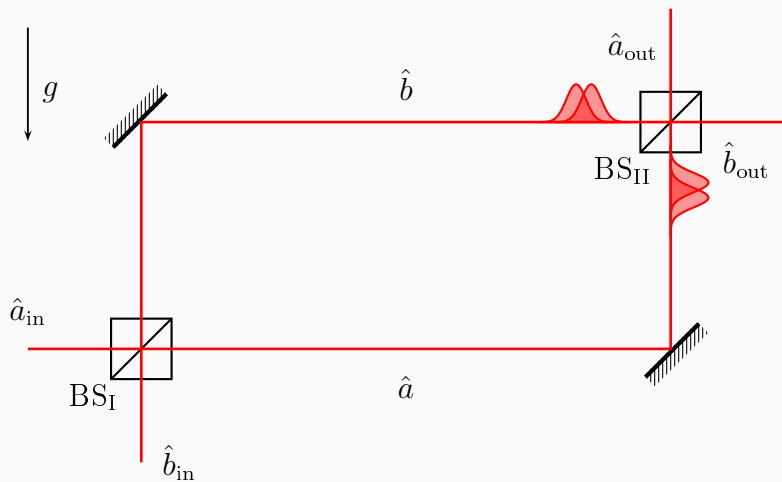
Photon Pairs



Photon Pairs



Photon Pairs



Static space-time metric

$$\mathbf{g} = -c^2 N(\mathbf{x})^2 dt^2 + h_{ij}(\mathbf{x}) dx^i dx^j$$

Riemann curvature

$$\mathbf{R}_{0i0j} = c^2 N \nabla_i \nabla_j N$$

$$\mathbf{R}_{ijkl} = R_{ijkl}$$

Gordon's optical metric

$$\tilde{\mathbf{g}} = -c^2 (N(\mathbf{x})/n(\mathbf{x}))^2 dt^2 + h_{ij}(\mathbf{x}) dx^i dx^j$$

Lagrangian

$$L[F] = -\frac{1}{4} \tilde{\mathbf{g}}^{ab} \tilde{\mathbf{g}}^{cd} F_{ac} F_{bd}$$

$$L[\varphi] = -\frac{1}{2} \tilde{\mathbf{g}}^{ab} (\nabla_a \varphi) (\nabla_b \varphi)$$

Neglecting polarisation degrees of freedom, consider scalar quantum field operator

$$\hat{\varphi}(t, \mathbf{x}) = \int \frac{d\omega}{\sqrt{2\omega}} [\hat{a}(\omega) e^{-i\omega t} u_{\omega}(\mathbf{x}) + \hat{a}^{\dagger}(\omega) e^{+i\omega t} \bar{u}_{\omega}(\mathbf{x})]$$

Mode functions satisfy

$$(n/N)\Delta(u_{\omega}N/n) + (n\omega/cN)^2 u_{\omega} = 0$$

Normalisation gives standard commutation relation

$$[a(\omega), a^{\dagger}(\omega')] = \delta(\omega - \omega')$$

Vacuum state $|0\rangle$

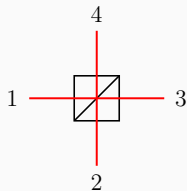
$$\hat{a}(\omega) |0\rangle = 0$$

Phases and Beam Splitters

First approximation: eikonal equation

$$\phi(\mathbf{x}_B) = \phi(\mathbf{x}_A) + \int_{\gamma} \frac{n\omega}{cN} \hat{k}_i dx^i$$

Beam splitter coupling at \mathbf{x}_* :



$$\hat{\varphi}_3(t, \mathbf{x}_*) = \mathcal{T}\hat{\varphi}_1(t, \mathbf{x}_*) + \mathcal{R}\hat{\varphi}_2(t, \mathbf{x}_*)$$

$$\hat{\varphi}_4(t, \mathbf{x}_*) = \mathcal{R}\hat{\varphi}_1(t, \mathbf{x}_*) + \mathcal{T}\hat{\varphi}_2(t, \mathbf{x}_*)$$

$$|\mathcal{T}|^2 + |\mathcal{R}|^2 = 1 \quad \bar{\mathcal{T}}\mathcal{R} + \bar{\mathcal{R}}\mathcal{T} = 0$$

$$\begin{pmatrix} \hat{a}_1^\dagger(\omega)\bar{u}_{1,\omega}(\mathbf{x}_*) \\ \hat{a}_2^\dagger(\omega)\bar{u}_{2,\omega}(\mathbf{x}_*) \end{pmatrix} = \begin{pmatrix} \mathcal{T} & \mathcal{R} \\ \mathcal{R} & \mathcal{T} \end{pmatrix} \begin{pmatrix} \hat{a}_3^\dagger(\omega)\bar{u}_{3,\omega}(\mathbf{x}_*) \\ \hat{a}_4^\dagger(\omega)\bar{u}_{4,\omega}(\mathbf{x}_*) \end{pmatrix}$$

The mode-couplings can be realised by unitary transformations \hat{U}_{BS} and \hat{U}_ϕ :

$$\hat{U}_{\text{BS}} |0\rangle = |0\rangle$$

$$\hat{U}_\phi |0\rangle = |0\rangle$$

$$\hat{U}_{\text{BS}} \hat{a}^\dagger \hat{U}_{\text{BS}}^\dagger = \frac{1}{\sqrt{2}}(a^\dagger + ib^\dagger) \quad \begin{array}{l} \mathcal{T} = 1/\sqrt{2} \\ \mathcal{R} = i/\sqrt{2} \end{array} \quad \hat{U}_{\text{BS}} \hat{b}^\dagger \hat{U}_{\text{BS}}^\dagger = \frac{1}{\sqrt{2}}(\hat{b}^\dagger + i\hat{a}^\dagger)$$

$$\hat{U}_\phi \hat{a}^\dagger \hat{U}_\phi^\dagger = a^\dagger e^{-i\phi'}$$

$$\hat{U}_\phi \hat{b}^\dagger \hat{U}_\phi^\dagger = \hat{b}^\dagger e^{-i\phi''}$$

Mach-Zehnder Interferometer

Consider single modes and symmetric beam splitters

$$|\psi\rangle = |1, 1\rangle = a^\dagger b^\dagger |0\rangle$$

$$\hat{U}_{\text{BS}} |\psi\rangle = \frac{i}{\sqrt{2}}(a^\dagger a^\dagger + b^\dagger b^\dagger) |0\rangle = \frac{i}{\sqrt{2}}(|2, 0\rangle + |0, 2\rangle)$$

$$\hat{U}_\phi \hat{U}_{\text{BS}} |\psi\rangle = \frac{i}{\sqrt{2}}(e^{-2i\phi'} |2, 0\rangle + e^{-2i\phi''} |0, 2\rangle)$$

$$\hat{U}_{\text{BS}} \hat{U}_\phi \hat{U}_{\text{BS}} |\psi\rangle = \frac{i}{2\sqrt{2}}(e^{-2i\phi'} - e^{-2i\phi''})(|2, 0\rangle - |0, 2\rangle) - \frac{1}{2}(e^{-2i\phi'} + e^{-2i\phi''}) |1, 1\rangle$$

Probability of finding both photons in the same mode

$$p = 2 \left| \frac{i}{2\sqrt{2}}(e^{2i\phi'} - e^{2i\phi''}) \right|^2 = \frac{1}{2}(1 - \cos(2\Delta\phi))$$

Compared to coherent states or single photons: **doubled** fringe frequency

The intermediate state $\frac{1}{\sqrt{2}}(|2, 0\rangle + |0, 2\rangle)$ is a special case of a maximally path-entangled “NOON” state $\frac{1}{\sqrt{2}}(|N, 0\rangle + |0, N\rangle)$.

Sensitivity to Curvature

Previous proposals for NOON-state interferometry: redshift (Anastopoulos & Hu 2021), frame dragging (Brady & Haldar 2021), PPN parameters (Rivera-Tapia *et al.* 2021) assumed linear dependence of $\Delta\phi \propto \Delta h$.

Linear dependence $\Delta N \propto \Delta h$ is explicable using the weak equivalence principle: $N = 1 + gz/c^2$ and $h_{ij} = \delta_{ij}$ is flat space-time.

Excess probability

$$\delta p = p - \tilde{p}$$

$-GM_{\oplus}/(R_{\oplus}+h)$
potential

const. + gh
potential

Sensitivity to Curvature

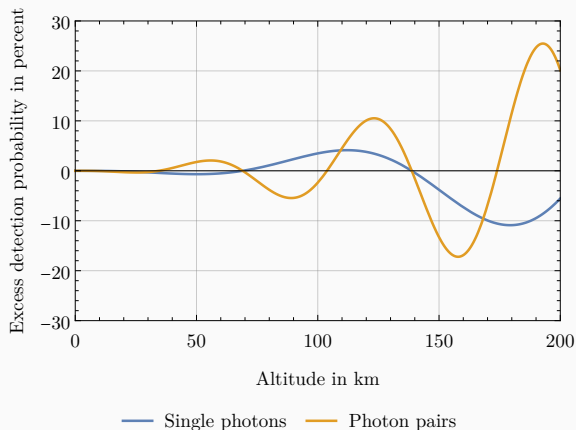


Figure 1: Excess probability, δp , for single photons and NOON states with wavelength $\lambda = 1500$ nm, spectral with $\delta\lambda = 1$ nm, and arm length $l = 100$ km.

Optical NOON-state interferometry experiments at satellite altitudes ...

- are beyond a Newtonian description of gravity
- are sensitive to gravity gradients
- simultaneously demonstrate signatures of quantum physics and general relativity

Anastopoulos, C. & Hu, B.-L. *Relativistic Particle Motion and Quantum Optics in a Weak Gravitational Field*. June 2021. arXiv: 2106.12514 [quant-ph].

Brady, A. J. & Haldar, S. Frame dragging and the Hong-Ou-Mandel dip: Gravitational effects in multiphoton interference. *Physical Review Research* **3**, 023024. doi:10.1103/PhysRevResearch.3.023024. arXiv: 2006.04221 [quant-ph] (Apr. 2021).

Colella, R., Overhauser, A. W. & Werner, S. A. Observation of Gravitationally Induced Quantum Interference. *Phys. Rev. Lett.* **34**, 1472–1474. doi:10.1103/PhysRevLett.34.1472 (June 1975).

Hilweg, C. *et al.* Gravitationally induced phase shift on a single photon. *New Journal of Physics* **19**, 033028. doi:10.1088/1367-2630/aa638f. arXiv: 1612.03612 [quant-ph] (Mar. 2017).

Kasevich, M. & Chu, S. Measurement of the gravitational acceleration of an atom with a light-pulse atom interferometer. *Applied Physics B: Lasers and Optics* **54**, 321–332. doi:10.1007/BF00325375 (May 1992).

Rivera-Tapia, M., Yáñez Reyes, M. I., Delgado, A. & Rubilar, G. *Outperforming classical estimation of Post-Newtonian parameters of Earth's gravitational field using quantum metrology*. Jan. 2021. arXiv: 2101.12126 [gr-qc].

Sorrentino, F. *et al.* Simultaneous measurement of gravity acceleration and gravity gradient with an atom interferometer. *Applied Physics Letters* **101**, 114106. doi:10.1063/1.4751112 (Sept. 2012).

Zych, M., Costa, F., Pikovski, I., Ralph, T. C. & Brukner, Č. General relativistic effects in quantum interference of photons. *Classical and Quantum Gravity* **29**, 224010. doi:10.1088/0264-9381/29/22/224010. arXiv: 1206.0965 [quant-ph] (Nov. 2012).